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Title: Nuclear Weapons Design Codes: A 50 Year Perspective

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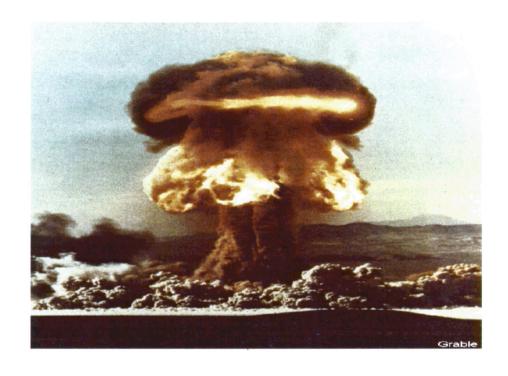


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Abstract: Nuclear Weapons Design Codes: A 50 Year Perspective

From Von Neumann using the Eniac to today's ASCI code developers using tens of teraflops on the ASCI computers; enormous progress has been made in modeling and understanding the physics and performance of nuclear explosives. And yet – we still have a long way to go if we truly expect to use our codes for predictive capability instead of relying on nuclear tests. This talk will review the significant hardware and software developments of the past decades and its impact on the development of nuclear design codes. There have been successes and failures, and despite the billions of dollars spent, significant challenges remain. How many flops and how many processors will it take to answer the questions? This talk will present the author's perspective on the problem, dig into a little history and try to speculate on the future.

Nuclear Weapons Design Codes: A 50 Year Perspective



NTS Grable 25 May 1953 15 kT Airburst

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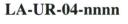
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Abstract

TITLE: Nuclear Weapons Design Codes: A 50 Year Perspective

From Von Neumann using the Eniac to today's ASCI code developers using tens of teraflops on the ASCI computers; enormous progress has been made in modeling and understanding the physics and performance of nuclear explosives. And yet – we still have a long way to go if we truly expect to use our codes for predictive capability instead of relying on nuclear tests. This talk will review the significant hardware and software developments of the past decades and its impact on the development of nuclear design codes. There have been successes and failures, and despite the billions of dollars spent, significant challenges remain. How many flops and how many processors will it take to answer the questions? This talk will present the author's perspective on the problem, dig into a little history and try to speculate on the future.







Outline

- What is THE application?
 - Answer: weapons design codes
- The Challenges:
 - Physics
 - Numerics
 - Computational
- Some History
- The Future







Why are we here?

Who is paying the bill?

LANL, LLNL, SNL

What is their primary mission?

Nuclear weapons

What is the purpose of ASCi?

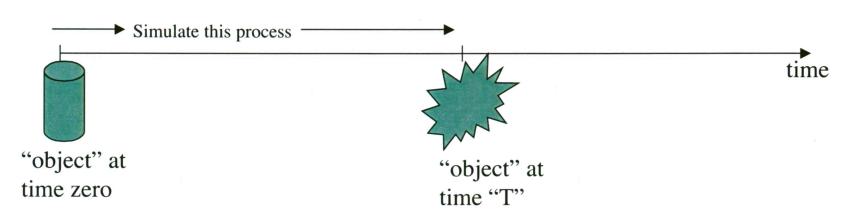
Predictive capability - nuclear weapons design codes







So...what is a design code?



- Model the <u>evolution</u> of the "object" from t=0 to t=T. Determine nuclear energy created and characteristics of the energy.
- Solve time dependent, non-linear, highly coupled, multi-physics, multi-dimensional differential equations describing "object" between t=0 and t=T.
- And get it right...and the stakes are HUGE...

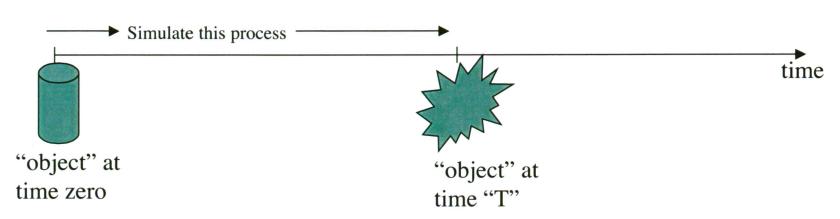
And....you currently cannot test to see if you've got it right!







Weapons Simulation: The Physics Challenge



Physics considerations

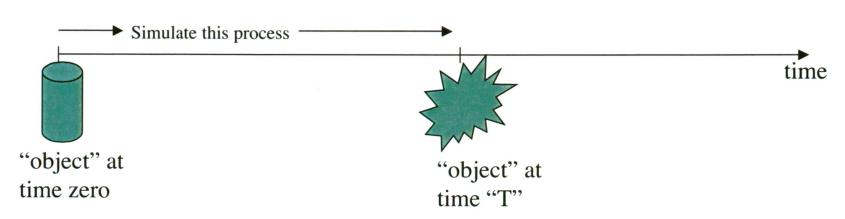
- 1. Assembly of critical mass hydrodynamics
- 2. Materials (nuclear fuels) fission and/or fusion (they "burn")
- 3. Materials go from room temperature (solids) to extremely high temps...rapidly
- 4. Interaction of nuclear products and energy with materials allows "object" to burn efficiently (or to not burn efficiently ie., it doesn't work)
- 5. Fundamental physical data over temperature and density space (cross-sections, material properties, etc.) is an enormous challenge
- 6. Effects of aging (What are the physical properties of the parts after 50+ years?)







Weapons Simulation: The Numerical Challenge



Numerical considerations:

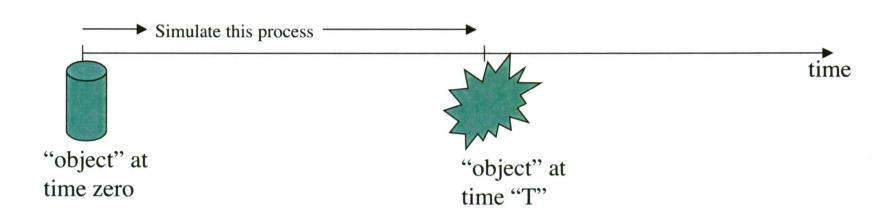
- 1. Complex time dependent, non-linear, coupled differential equations
- 2. 3D
- 3. Enormously large matrices to be "solved"
- 4. Quantify uncertainty without experiments







Weapons Simulation: The Computational Challenge



- 1. Huge codes:
 - a. 0.5-1.0 million lines of code, massively parallel
 - b. Complex SQA problems
 - c. Code components in multiple languages
- 2. Really BIG massively parallel computers (tens/hundreds of tera-ops)
- 3. Massive I/O (data transfer) challenges
- 4. Visualization challenges
- 5. Language issues
- 6. Cosmic radiation (keep the nodes running)







Hydrodynamics: the momentum equation

$$(F = MA)$$

$$\frac{d\vec{v}}{dt} = -\frac{1}{\rho}\vec{\nabla}p$$

Differential eqn => difference equation:

How many spatial cells to resolve? 1000/D? => 1000**3 How many time steps to resolve? 10,000?

10**13 dp*

That's a pretty BIG problem! A minimum 3D problem needs ~ 10**10 dp

The historic solution....reduce the dp's to fit the computer and your lifestyle.

* dp = discretization points







Neutron transport: the Boltzmann equation

$$\frac{1}{v(E)} \frac{\partial \psi(\vec{r}, E, \underline{\Omega}, t)}{\partial t} + \underline{\Omega} \cdot \nabla \psi(\vec{r}, E, \underline{\Omega}, t) + \sigma_t(\vec{r}, E, t) \psi(\vec{r}, E, \underline{\Omega}, t) =$$

$$\int_{4\pi^0}^{\infty} \sigma_s(\vec{r}, E' \to E, \underline{\Omega}' \to \underline{\Omega}, t) \psi(\vec{r}, E', \underline{\Omega}', t) dE' d\Omega' +$$

$$\frac{\chi(E)}{4\pi} \int_{0}^{\infty} v(\vec{r}, E', t) \sigma_f(\vec{r}, E', t) \phi(\vec{r}, E', t) dE' + S_{ext}(\vec{r}, E, \underline{\Omega}, t)$$

Differential eqn => difference equation:

How many spatial cells to resolve? 1000/D? => 1000**3? How many energy groups to resolve? 100? How many angular groups to resolve? 100? How many time steps to resolve? 10,000?



That's a really BIG problem!

The historic solution...reduce the dp's to fit the computer and your lifestyle.







The Transport Equation

There is a whole lot more to solving the transport equation than dp's! (re: Jim Morel's talk)

- Books have been written on the subject and many careers have been dedicated to the subtleties of the numerics particularly when the dp's are "few" and the dp spacing is complex.
- Matrix solvers, pre-conditioners and acceleration schemes are critical.
- Getting the limits correct...through thick and thin.







Some History....the early years

- 1951 the ENIAC (Von Neumann, LANL) (10**4 dp):
 - the first design code
 - 50 cells, 100 time steps, clock time 6 months
 - language HPT (holes in paper tape)
- The 60's IBM 709,7090, STRETCH (10**6 dp):
 - 10**3 cells, 10**3 time steps, clock time 10's of hours
 - assembly language
- 1968 CDC6600 (10**7 dp):
 - Von Neumann's "code"
 - 600 cells, 10**4 time steps, clock time 100's of hours
 - FORTRAN
- 1970 the CDC STAR (LLNL):
 - vectorization but not productive







Some History....the mid years

- The early 70's The CDC 7600 (10**8 dp):
 - 10**4 cells, 10**4 time steps, clock time 10's of hours,
- The mid 70's CRAY 1 (10**9 dp):
 - vectorization works!
 - 10**5 cells, 10**4 time steps, clock time 10's of hours,
- 1981
 - Salishan #1
 - WPC "designs" parallel CRAY architecture for Monte Carlo!
 - Vectorization of Monte Carlo was difficult (but doable) => Bill Martin, Forrest Brown
 - Parallel architecture seems a natural for Monte Carlo
- The late 80's Cray XMP, YMP (10**10 dp):
 - Parallel architecture
 - 10**5 cells, 10**4 time steps, 10's of hours
 - but codes still mostly serial







Some History....the recent years

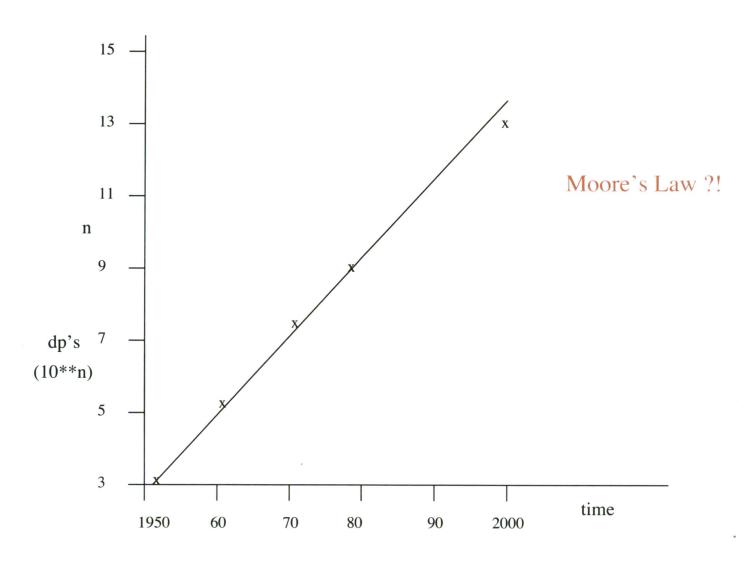
- The early 90's Connection Machine (LANL), the Meiko (LLNL) and others:
 - Nuclear testing ends
 - No vectorization
 - some venturing into parallel processing
 - Not heavily used by design codes
 - Code porting issues
- The mid 90's DEC Alpha Clusters (LLNL), UNIX workstations:
 - vectorization goes away
 - codes ported to UNIX workstations (the end of LRLTRAN, LTSS)
 - C, C++ for a few but FORTRAN for most
 - Some parallelism, but minimal
- The late 90's ASCi Blue (SGI, IBM) (10**11 dp):
 - Lots of \$\$\$ being spent
 - Emphasis on 3D
 - 10**6 cells, 10**4 time steps, clock time 100's hours
- The early 2000's ASCi HP-Compaq Q (10**13 dp):
 - 10**8 cells, 10**4 time steps, clock time 100's hours







Dp's vs time











What do we need (architecture)? (1)

- Fast CPUs
 - not all problems are communications bound
- Lots of memory
 - would like a homogenous resource
 - some problems using 2GB/PE
- High speed interconnect network
 - technology that maintains processor/network speed ratio of 10-20 Flop/MBs
- Truly global parallel file system
 - data on PFS to be visible from any processing element in the computing center







What do we need (architecture)? (2)

- High bandwidth secondary storage (including the supporting network)
 - How do we efficiently save data without slowing down on-going processing?
 - generating up to 200 GB per restart dump (every 30-60 minutes).
- High Capacity Visualization
 - tied to data movement issues, real-time processing, and data reduction
 - generating 140 GB of data per time slice
- Stability
 - successfully port and run on system N+1 before the system N+2 arrives







What have we learned?

- We will use all the capacity that is provided.
- Some of our applications scale well to 1000's of PEs.
 - Two of our main ASC codes scale optimally to at least 1800 PEs.
- We are a flexible and inventive bunch.
 - We will overcome or at least minimize shortcomings in any architecture.
 - We waste a lot of time overcoming.







Robustness issues

- User frustration issue...efficiency factor ~.9 (<100 PEs) to ~.3 (>1000 PEs)
- Frequent restart dumps and automated submission/monitoring scripts ease the pain.
- Example: on Q series ASCi machines, Mean Time To Interrupt ~ 6 hours.
 - MTTI might become more of an issue as we scale to PetaFlop processing using a cluster approach with many more parts that can randomly fail.







User experiences

- Turn around for many problems less than a day.
- Turn around for "hero" problems days to month(s).
- Typically months of debugging and trial runs <u>prior</u> to the "successful" hero run.
- The goal: Provide answers while we still remember what we're looking for. (<50 hrs cpu ~ a week user time)







The Typical Code

- 0.5 1.0 Million lines
- Spatial discretization may be extremely complex
- Fortran 77 95, C, C++
- 20-35 year life cycle
 - Moderate rewrites every 10-15 years
- MPI, maybe some threading
- Totalview debugger (and "print")







The End





